

403(2): Precession from the ECE2 force equation of orbits

Consider the ECE2 force equation of orbits:

$$\underline{F} = m \underline{\ddot{r}} = -\underline{\nabla} \phi + \underline{\omega} \phi \quad (1)$$

where

$$\phi = -\frac{m\Gamma}{r} \quad (2)$$

$\Gamma$  is the gravitational potential and  $\underline{\omega}$  is the spin connection.  
In general, in plane polar coordinates  $(r, \phi)$ :

$$\underline{\omega} = \omega_r \underline{e}_r + \omega_\phi \underline{e}_\phi \quad (3)$$

so eq. (1) becomes:

$$\ddot{r} - r\dot{\phi}^2 = -\frac{m\Gamma}{r^2} - \frac{m\Gamma}{r} \omega_r \quad (4)$$

and

$$r\ddot{\phi} + 2\dot{r}\dot{\phi} = -\frac{m\Gamma}{r} \omega_\phi \quad (5)$$

from eq. (4):

$$F = -\frac{m\Gamma}{r^2} (1 + \omega_r r) \quad (6)$$

the vacuum produces the force:

$$F(\text{vac}) = -\frac{m\Gamma \omega_r}{r} = -\frac{\partial U(r)}{\partial r} \quad (7)$$

so

$$U(\text{vac}) = m\Gamma \int \frac{\omega_r}{r} dr \quad (8)$$

The Lagrangian is:

$$L = T - U \quad (9)$$

and the Hamiltonian is:

$$H = T + U \quad (10)$$

where

$$T = \frac{1}{2} m (\dot{r}^2 + r^2 \dot{\phi}^2) \quad (11)$$

$T$  is kinetic energy. The Euler Lagrange equation:

$$2) \quad \frac{\partial \mathcal{L}}{\partial r} = \frac{d}{dt} \left( \frac{\partial \mathcal{L}}{\partial \dot{r}} \right) \quad - (12)$$

gives equation (4), and

$$\frac{\partial \mathcal{L}}{\partial \phi} = \frac{d}{dt} \left( \frac{\partial \mathcal{L}}{\partial \dot{\phi}} \right) \quad - (13)$$

gives the angular momentum:

$$L = \frac{\partial \mathcal{L}}{\partial \dot{\phi}} = m r^2 \dot{\phi} \quad - (14)$$

with

$$\frac{dL}{dt} = 0 \quad - (15)$$

So  $L$  is a constant of motion. The other constant of motion is the Hamiltonian (10), so:

$$\frac{dH}{dt} = 0 \quad - (16)$$

In Q above,  $U$  is the total potential energy:

$$U = -\frac{mMg}{r} + mg \int \frac{\omega_r}{r} dr \quad - (17)$$

As in UFT 402, eq. (14) gives:

$$r\ddot{\phi} + 2\dot{r}\dot{\phi} = 0 \quad - (18)$$

so

$$\omega_{\phi} = 0 \quad - (19)$$

The equation (4) can be transformed into the generalized Binet equation:

$$\frac{d^2}{d\phi^2} \left( \frac{1}{r} \right) + \frac{1}{r} = \frac{1}{\alpha} (1 + r\omega_r) \quad - (20)$$

which is the vacuum corrected Binet equation.

3) In the limit  $\omega_r \rightarrow 0$  - (21)

The solution of eq. (20) is the well known circular section:

$$r = \frac{d}{1 + \epsilon \cos \phi} \quad - (22)$$

Now follow the method of note 403(1) to obtain the precession of the half right orbit. In the limit (21) the half right orbit is:

$$d = r \quad - (23)$$

and is defined by

$$\cos \phi = 0 \quad \phi = \frac{\pi}{2} \quad - (24)$$

For the circular section (22), the half right orbit is defined by

$$\frac{d^2}{dr^2} = 0 \quad - (24)$$

so eq. (23) results from the Binet equation:

$$\frac{d^2}{dr^2} + \frac{1}{r} = \frac{1}{d} \quad - (25)$$

As in note 403(1) it is assumed that eq. (24) is approximately true for eq. (20), so:

$$\frac{1}{r} \sim \frac{1}{d} (1 + r\omega_r) \quad - (26)$$

Therefore  $d/r$  has increased from 1 at the half right orbit to  $1 + r\omega_r$ . If the spin correction component  $\omega_r$  is negative valued,  $d/r$  decreases from 1 to  $1 - r|\omega_r|$ .

The ellipse precesses as shown in Fig (1):

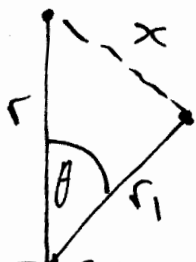


Fig (1)

At the half right rule:

$$r = d \quad - (27)$$

$$r_1 = d(1 - d|\omega_r|) \quad - (28)$$

By the triangle rule, as in Fig (1):

$$x^2 = r^2 + r_1^2 - 2rr_1 \cos \theta \quad - (29)$$

So

$$r_1 = r \cos \theta \pm \left( r^2 \cos^2 \theta - (r^2 - x^2) \right)^{1/2}$$

$$\sim r \cos \theta \quad - (30)$$

for small precessions:

$$x \rightarrow 0, \cos \theta \rightarrow 1 \quad - (31)$$

so the precession at the half right rule is:

$$\cos \theta \sim \frac{r_1}{r} \quad - (32)$$

i.e.

$$\cos \theta \sim 1 - d|\omega_r| \quad - (33)$$

Finally use:

$$\cos \theta = 1 - \frac{\theta^2}{2!} + \dots \quad - (34)$$

to find that

$$\theta^2 \sim 2d|\omega_r| \quad - (35)$$

Exact agreement with experimental data is found by adjusting  $|\omega_r|$ .