

366 (1) : Self Consistency Check, Newton's Field

Given:

$$x = \frac{d}{\left(\frac{2}{r} - \frac{1}{a}\right)^{3/2}} \left(\frac{\epsilon}{d} \sin^2 \theta + \frac{1}{r} \cos \theta \right) \quad - (1)$$

$$y = \frac{d}{\left(\frac{2}{r} - \frac{1}{a}\right)^{3/2}} \left(\frac{1}{r} \sin \theta - \frac{\epsilon \sin \theta \cos \theta}{d} \right) \quad - (2)$$

it must be shown that

$$\frac{1}{(x^2 + y^2)^{1/2}} = \frac{2}{r} - \frac{1}{a} \quad - (3)$$

From eqs (1) and (2):

$$\begin{aligned} x^2 + y^2 &= \frac{d}{\left(\frac{2}{r} - \frac{1}{a}\right)^3} \left(\left(\frac{\epsilon}{d} \sin^2 \theta + \frac{1}{r} \cos \theta \right)^2 + \left(\frac{1}{r} \sin \theta - \frac{\epsilon \sin \theta \cos \theta}{d} \right)^2 \right) \\ &= \frac{d}{\left(\frac{2}{r} - \frac{1}{a}\right)^3} \left(\left(\frac{\epsilon}{d} \right)^2 \sin^4 \theta + \frac{2\epsilon}{dr} \sin^2 \theta \cos \theta + \frac{1}{r^2} \cos^2 \theta \right. \\ &\quad \left. + \frac{1}{r^2} \sin^2 \theta - \frac{2\epsilon}{rd} \sin^2 \theta \cos \theta + \left(\frac{\epsilon}{d} \right)^2 \sin^2 \theta \cos^2 \theta \right) \end{aligned}$$

$$= \frac{d}{\left(\frac{2}{r} - \frac{1}{a}\right)^3} \left(\frac{1}{r^2} + \left(\frac{\epsilon}{d} \right)^2 \sin^2 \theta \right) \quad - (4)$$

$$\text{Now use } \frac{1}{r^2} = \frac{1}{d^2} (1 + \epsilon \cos \theta)^2 \quad - (5)$$

So:

$$\begin{aligned}x^2 + y^2 &= \frac{\alpha}{\left(\frac{2}{r} - \frac{1}{a}\right)^3} \left(1 + 2\epsilon \cos\theta + \epsilon^2\right) \\&= \frac{1}{\alpha \left(\frac{2}{r} - \frac{1}{a}\right)^3} \left(1 + 2\left(\frac{\alpha}{r} - 1\right) + \epsilon^2\right) \\&= \left(\frac{2}{r} + \frac{\epsilon^2 - 1}{\alpha}\right) / \left(\frac{2}{r} - \frac{1}{a}\right)^3 \\&= \left(\frac{2}{r} - \frac{1}{a}\right) / \left(\frac{2}{r} - \frac{1}{a}\right)^3 \\&= \frac{1}{\left(\frac{2}{r} - \frac{1}{a}\right)^2} \quad - (6)\end{aligned}$$

So

$$\frac{1}{(x^2 + y^2)^{1/2}} = \frac{2}{r} - \frac{1}{a} \quad - (7)$$

A. E. D.

In general, for any planar orbit:

$$\underline{g} = (\underline{v} \cdot \underline{\nabla}) \underline{v} \quad - (8)$$

also

$$\underline{v} = \dot{r} \underline{e}_r + r\dot{\theta} \underline{e}_\theta \quad - (9)$$

The Newtonian orbit is:

$$r = \frac{\alpha}{1 + \epsilon \cos\theta} \quad - (10)$$

), but the experimentally observed orbit is a precessing ellipse. For the precessing ellipse:

$$\underline{v} = (MG)^{1/2} \frac{(-\underline{y}_i + \underline{x}_j)}{(x^2 + y^2)^{3/4}} \quad (11)$$

so x and y must be found for the precessing ellipse terms of r and θ . Having done this, the force law for the observed orbit can be found using eq. (8):

$$\underline{F} = m\underline{g} = m(\underline{v} \cdot \nabla) \underline{v} \quad (12)$$

where \underline{v} is the orbital velocity observed experimentally.

It is known experimentally that the perihelion is displaced in every revolution by:

$$\Delta = \frac{6\pi MG}{ac^2(1-e^2)} \quad (13)$$

i. e. $\theta = 2\pi \rightarrow 2\pi \left(1 + \frac{3MG}{ac^2(1-e^2)} \right) \quad (14)$

$= x\theta$

$$\boxed{rc = \frac{1 + \frac{3MG}{ac^2(1-e^2)}}{1 + e \cos(x\theta)}}; \text{ so } \boxed{r = \frac{d}{1 + e \cos(x\theta)}} \quad (15)$$